

Instantons in $SU(N)$ (pure) Yang-Mills Theory

Would like to calculate the overlap between equivalent vacua of winding numbers differing by 1: $\Delta n = 1$

$$\langle n+1 | \hat{H} | n \rangle$$

We look for the instanton solution to the Euclideanized Yang-Mills action:

$$S_E[A] = \int d^4 x_E \frac{+1}{4} (F_{\mu\nu}^a F_{\mu\nu}^a)_E = \int d^4 x_E \frac{+1}{2} \left((\vec{E}_E^a)^2 + (\vec{B}_E^a)^2 \right)$$

(strictly positive)

That is, we aim to extremize the action subject to the boundary condition:

System starts at n :

$$A_\mu^{\text{inst}}(\tau \rightarrow -\infty) = \frac{-i}{g} U^{(n)}(\vec{x}) \partial_\mu U^{\dagger(n)}(\vec{x})$$

↑
Eucl. time

System ends at $n+1$:

$$A_\mu^{\text{inst}}(\tau \rightarrow +\infty) = \frac{-i}{g} U^{(n+1)}(\vec{x}) \partial_\mu U^{\dagger(n+1)}(\vec{x})$$

Pure gauges

AND

$$A_\mu^{\text{inst}}(\vec{x} \rightarrow \infty) = 0 \quad \Rightarrow \quad \partial_\mu U(\vec{x}) = 0$$

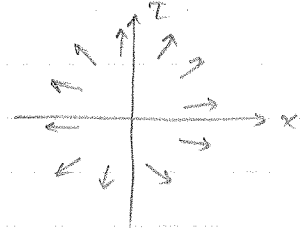
i.e. the $U(\vec{x})$ approaches a constant as $\vec{x} \rightarrow \infty$.

These boundary conditions can more easily be satisfied if we consider (hyper) spherically symmetric field configurations, such that gauge fields on the boundary 3-sphere have winding number = 1.

As with other large field configuration (topological) solutions, set angles on the hypersphere equal to angles on the vacuum manifold:

$$\text{Let } U(x_\mu) = a_0(x_\mu) \mathbb{1}_{2 \times 2} + i \vec{a}(x_\mu) \cdot \vec{\sigma}$$

$$\text{subject to } \sum_{\mu=0}^3 a_\mu a_\mu = 1$$



Take $a_\mu = \hat{x}_\mu$, ← unit 4-vector

$$\text{so that } U(\hat{x}_\mu) = \sum_{\mu} \hat{x}_\mu \sigma_\mu, \quad \sigma_\mu = (1, i\vec{\sigma}) \leftarrow \text{obey: } \sigma_\mu^\dagger \sigma_\nu + \sigma_\nu^\dagger \sigma_\mu = 2\delta_{\mu\nu}$$

Hence, the gauge fields on the hypersphere at infinity is:

$$\begin{aligned} A_\mu(|x_\mu| \rightarrow \infty) &= \frac{-i}{g} U(\hat{x}) \partial_\mu U^\dagger(\hat{x}) \\ &= \frac{-i}{g} \hat{x}_\nu \sigma_\nu \partial_\mu (\hat{x}_\rho \sigma_\rho^\dagger) \\ &= \frac{-i}{g} \hat{x}_\nu \sigma_\nu \left(\frac{\delta_{\mu\rho} - \hat{x}_\mu \hat{x}_\rho}{|x|} \right) \sigma_\rho^\dagger \quad \left\{ \text{use } \partial_\mu \hat{x}_\rho = \frac{\delta_{\mu\rho} - \hat{x}_\mu \hat{x}_\rho}{|x|} \right. \\ &= \frac{-i}{g} \underbrace{\sigma_\nu \sigma_\rho^\dagger}_{\uparrow} \hat{x}_\nu \left(\frac{\delta_{\mu\rho} - \hat{x}_\mu \hat{x}_\rho}{|x|} \right) \end{aligned}$$

$$\text{Now: } \sigma_\mu \sigma_\nu^\dagger = \begin{pmatrix} \mathbb{1} = \delta_{00} \mathbb{1} & -i\vec{\sigma}_j \\ i\vec{\sigma}_i & \sigma_i \sigma_j \end{pmatrix}$$

Note: $\sigma_\nu^\dagger = (1, -i\vec{\sigma})$, $\vec{\sigma}$ hermitian

$$\uparrow \text{ write: } \sigma_i \sigma_j = \frac{1}{2} \{ \sigma_i, \sigma_j \} + \frac{1}{2} [\sigma_i, \sigma_j]$$

$$= \frac{1}{2} 2\delta_{ij} + \frac{1}{2} 2i\epsilon^a_{ij} \sigma^a$$

$$= \delta_{ij} + i\epsilon^a_{ij} \sigma^a$$

$$\text{so, } \sigma_\mu \sigma_\nu^\dagger = \begin{pmatrix} \delta_{00} & | & 0 \\ \hline 0 & | & \delta_{ij} \end{pmatrix} + i \begin{pmatrix} 0 & | & -\delta_{\mu 0} \delta_\nu^a \\ \hline \delta_{\nu 0} \delta_\mu^a & | & \epsilon^a_{ij} \end{pmatrix} \sigma^a$$

$a = \{1, 2, 3\}$
 $\mu, \nu = \{0, 1, 2, 3\}$

$$= \delta_{\mu\nu} + i \underbrace{(\epsilon^a_{\mu\nu} + \delta_\mu^a \delta_{\nu 0} - \delta_\nu^a \delta_{\mu 0})}_{\text{= } \epsilon^a_{\mu\nu} \text{ (if } \mu, \nu \neq 0)} \sigma^a = \delta_{\mu\nu} + i\eta^a_{\mu\nu} \sigma^a$$

4 Hooft symbols Maps $SO(3) \times SO(3) \rightarrow SO(4)$

Defⁿ $\eta^a_{\mu\nu} = \epsilon^a_{\mu\nu} + \delta^a_\mu \delta_{\nu 0} - \delta^a_\nu \delta_{\mu 0}$

There are three ($a = \{1, 2, 3\}$)
4x4 matrices ($\mu, \nu = \{0, 1, 2, 3\}$)

$\epsilon^a_{\mu\nu} = 0$ if $\{\mu=0 \text{ or } \nu=0\}$ $\delta^a_\mu = 0$ if $\{\mu \neq 0\}$ $\delta^a_\nu = 0$ if $\{\nu \neq 0\}$

Explicitly, $\eta^1_{\mu\nu} = \begin{pmatrix} 0 & -1 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 \\ 0 & -1 & 0 & 0 \end{pmatrix}$ $\eta^2_{\mu\nu} = \begin{pmatrix} 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \\ 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix}$ $\eta^3_{\mu\nu} = \begin{pmatrix} 0 & 0 & 0 & -1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ 1 & 0 & 0 & 0 \end{pmatrix}$

Properties:

$\eta^a_{\mu\nu} = \frac{1}{2} \hat{\epsilon}_{\mu\nu\rho\sigma} \eta^a_{\rho\sigma}$ (self-dual) $\hat{\epsilon}_{0123} = -1$
($\hat{\epsilon}_{1230} = +1$)

$\eta^a_{\mu\nu} = -\eta^a_{\nu\mu}$ (antisymmetric)

$\eta^a_{\mu\nu} \eta^b_{\mu\nu} = 4\delta^{ab}$

$\eta^a_{\mu\nu} \eta^a_{\mu\rho} = 3\delta_{\nu\rho}$

$\eta^a_{\mu\nu} \eta^a_{\mu\nu} = 12$

Proofs:
Leave as exercises

$\eta^a_{\mu\nu} \eta^b_{\mu\rho} = \delta^{ab} \delta_{\nu\rho} + \epsilon^{abc} \eta^c_{\nu\rho}$

Interesting to note:

If we let indices μ & ν to run $\{1, 2, 3, 4\}$ instead of $\{0, 1, 2, 3\}$
we would get:

$\eta^1 = \begin{pmatrix} & & 1 \\ & 1 & \\ -1 & & \\ & & \end{pmatrix}$ $\eta^2 = \begin{pmatrix} & -1 & \\ & & 1 \\ 1 & & \\ & -1 & \end{pmatrix}$ $\eta^3 = \begin{pmatrix} & & 1 \\ -1 & & \\ & & \\ & & 1 \\ & -1 & \end{pmatrix}$

These are precisely ($2i$ times) the generators of custodial weak isospin
in the Higgs sector of the SM:

$(T^a_{\text{cus}}) = \frac{1}{2i} (\eta^a)$

Solution to exercises:

$$\begin{aligned} \eta^a{}_{\mu\nu} \eta^a{}_{\mu\rho} &= (\epsilon^a{}_{\mu\nu} + \delta^a{}_{\mu} \delta_{\nu 0} - \delta^a{}_{\nu} \delta_{\mu 0}) (\epsilon^a{}_{\mu\rho} + \delta^a{}_{\mu} \delta_{\rho 0} - \delta^a{}_{\rho} \delta_{\mu 0}) \\ &\stackrel{①}{=} \epsilon^a{}_{\mu\nu} \epsilon^a{}_{\mu\rho} + \epsilon^a{}_{\mu\nu} \delta^a{}_{\mu} \delta_{\rho 0} - \epsilon^a{}_{\mu\nu} \delta^a{}_{\rho} \delta_{\mu 0} \\ &\quad + \delta^a{}_{\mu} \delta_{\nu 0} \epsilon^a{}_{\mu\rho} + \delta^a{}_{\mu} \delta_{\nu 0} \delta^a{}_{\mu} \delta_{\rho 0} - \delta^a{}_{\mu} \delta_{\nu 0} \delta^a{}_{\rho} \delta_{\mu 0} \\ &\quad - \delta^a{}_{\nu} \delta_{\mu 0} \epsilon^a{}_{\mu\rho} - \delta^a{}_{\nu} \delta_{\mu 0} \delta^a{}_{\mu} \delta_{\rho 0} + \delta^a{}_{\nu} \delta_{\mu 0} \delta^a{}_{\rho} \delta_{\mu 0} \end{aligned}$$

- ① $\epsilon^a{}_{\mu\nu} \epsilon^a{}_{\mu\rho} = 2(\delta_{\nu\rho} - \delta_{\nu 0} \delta_{\rho 0})$ ← like $\epsilon_{ijk} \epsilon_{ijl} = 2\delta_{kl}$, but must remove $\delta_{(v=0)(\rho=0)}$ contribution since $\epsilon^a{}_{\mu\nu} = 0$ for $\mu=0$ or $\nu=0$.
- ② $\delta^a{}_{\mu} \delta_{\nu 0} \delta^a{}_{\mu} \delta_{\rho 0} = 3 \delta_{\nu 0} \delta_{\rho 0}$
- ③ $\delta^a{}_{\nu} \delta_{\mu 0} \delta^a{}_{\rho} \delta_{\mu 0} = \delta^a{}_{\nu} \delta^a{}_{\rho} = \delta_{\nu\rho} - \delta_{\nu 0} \delta_{\rho 0}$ ← again, must remove $\delta_{(v=0)(\rho=0)}$ contribution.

S_g

$$\begin{aligned} \eta^a{}_{\mu\nu} \eta^a{}_{\mu\rho} &= 2(\delta_{\nu\rho} - \delta_{\nu 0} \delta_{\rho 0}) + 3 \delta_{\nu 0} \delta_{\rho 0} + \delta_{\nu\rho} - \delta_{\nu 0} \delta_{\rho 0} \\ &= 3 \delta_{\nu\rho} \quad \square \end{aligned}$$

Then, setting $\nu = \rho$ and summing gives $\delta_{pp} = 4$, so that

$$\eta^a{}_{\mu\nu} \eta^a{}_{\mu\nu} = 3 \times 4 = 12 \quad \square$$

So, in terms of 't Hooft symbols, we have at the surface,

$$\begin{aligned}
 A_\mu(|x| \rightarrow \infty) &= \frac{-i}{g} (\delta_{\nu\mu} + i\eta^a_{\nu\mu} \sigma^a) \hat{x}_\nu \left(\frac{\delta_{\mu\rho} - \hat{x}_\mu \hat{x}_\rho}{|x|} \right) \\
 &= \frac{-i}{g} \frac{1}{|x|} \left(\hat{x}_\mu - \hat{x}_\nu \hat{x}_\mu \hat{x}_\nu + i\eta^a_{\nu\mu} \hat{x}_\nu \sigma^a - i\eta^a_{\nu\mu} \hat{x}_\nu \hat{x}_\mu \hat{x}_\rho \sigma^a \right) \\
 &= \frac{-i}{g} \frac{1}{|x|} (-i\eta^a_{\mu\nu} \hat{x}_\nu \sigma^a) \\
 &\quad \text{write as } x_\nu/|x|
 \end{aligned}$$

$$A_\mu \equiv A_\mu^a(|x| \rightarrow \infty) \frac{\sigma^a}{2} = \frac{-1}{g} \frac{1}{x_E^2} \eta^a_{\mu\nu} x_\nu \sigma^a$$

Dropping the σ^a , we have

$$A_\mu^a(|x| \rightarrow \infty) = \frac{-2}{g} \frac{1}{x_E^2} \eta^a_{\mu\nu} x_\nu \quad \text{on the boundary 3-sphere.}$$

Using this result as a guide, we can write down an ansatz that would be valid for all (Euclidean) space-time — not just on the hypersphere:

Ansatz:

$$A_\mu^a(x) = \frac{-2}{g} \frac{f(x^2)}{x^2} \eta^a_{\mu\nu} x_\nu, \quad \begin{array}{l} f(x^2) \text{ is a function to be solved for} \\ \text{satisfying } f(\infty) \rightarrow 1 \\ f(0) \rightarrow 0 \text{ (for continuous solutions)} \end{array}$$

Construct the field-strength tensor:

$$F_{\mu\nu}^a = \overset{\textcircled{1}}{\partial_\mu A_\nu^a} - \overset{\textcircled{2}}{\partial_\nu A_\mu^a} - \overset{\textcircled{3}}{g f^{abc} A_\mu^b A_\nu^c}$$

$$\begin{aligned}
 \textcircled{1} \quad \partial_\mu A_\nu^a &= \frac{-2}{g} \eta^a_{\nu\rho} \partial_\mu \left(\frac{f(x^2) x_\rho}{x^2} \right) \quad \text{use quotient rule.} \\
 &= \frac{-2}{g} \eta^a_{\nu\rho} \frac{1}{x^4} \left[\left(f'(x^2) 2x_\mu x_\rho + f(x^2) \delta_{\mu\rho} \right) x^2 - f(x^2) x_\rho 2x_\mu \right]
 \end{aligned}$$

Factor out 2:

$$= \frac{-4}{g} \left[\frac{1}{x^4} \left(x^2 f'(x^2) - f(x^2) \right) \eta^a_{\nu\rho} x_\rho x_\mu + \frac{1}{x^2} f(x^2) \frac{1}{2} \eta^a_{\nu\mu} \right]$$

Similarly,

$$\textcircled{2} \partial_\nu A_\mu^a = \frac{-4}{g} \left[\frac{1}{x^4} (x^2 f'(x^2) - f(x^2)) \eta^a_{\mu\rho} x_\rho x_\nu + \frac{1}{x^2} f(x^2) \frac{1}{2} \eta^a_{\mu\nu} \right]$$

Then

$$\begin{aligned} \partial_\mu A_\nu^a - \partial_\nu A_\mu^a &= \frac{-4}{g} \left[\frac{1}{x^4} (x^2 f'(x^2) - f(x^2)) (\eta^a_{\nu\rho} x_\rho x_\mu - \eta^a_{\mu\rho} x_\rho x_\nu) \right. \\ &\quad \left. + \frac{1}{x^2} f(x^2) \frac{1}{2} (\eta^a_{\nu\mu} - \eta^a_{\mu\nu}) \right] \\ &= \frac{-4}{g} \left[\frac{1}{x^4} (x^2 f'(x^2) - f(x^2)) (\eta^a_{\nu\rho} x_\rho x_\mu - \eta^a_{\mu\rho} x_\rho x_\nu) - \frac{1}{x^2} f(x^2) \eta^a_{\mu\nu} \right] \end{aligned}$$

Now

$$\begin{aligned} \textcircled{3} -g f^{abc} A_\mu^b A_\nu^c &= g f^{abc} \left(\frac{-2}{g} \frac{f(x^2)}{x^2} \right)^2 \eta^b_{\mu\rho} \eta^c_{\nu\sigma} x_\rho x_\sigma \\ &= \frac{-4}{g} \frac{f^2(x^2)}{x^4} \underbrace{f^{abc} \eta^b_{\mu\rho} \eta^c_{\nu\sigma}} x_\rho x_\sigma \end{aligned}$$

Assume instanton solution in $SU(2)$ subgroup: $f^{abc} \rightarrow \epsilon^{abc}$

$$\begin{aligned} &= \frac{-4}{g} \frac{f^2(x^2)}{x^4} (\delta_{\mu\nu} \eta^a_{\rho\sigma} - \delta_{\mu\sigma} \eta^a_{\rho\nu} - \delta_{\rho\sigma} \eta^a_{\mu\nu} - \delta_{\rho\nu} \eta^a_{\mu\sigma}) x_\rho x_\sigma \\ &= \frac{-4}{g} \frac{f^2(x^2)}{x^4} \left(\underbrace{-\eta^a_{\rho\nu} x_\rho x_\mu}_{\substack{\uparrow \\ \sigma}} + \eta^a_{\mu\nu} x^2 - \underbrace{\eta^a_{\mu\sigma} x_\sigma x_\nu}_{\substack{\uparrow \\ \rho}} \right) \\ &= \frac{-4}{g} \left[\frac{-f^2(x^2)}{x^2} \eta^a_{\mu\nu} + \frac{1}{x^4} f^2(x^2) (\eta^a_{\nu\rho} x_\rho x_\mu - \eta^a_{\mu\rho} x_\rho x_\nu) \right] \end{aligned}$$

So, adding everything together,

$$\begin{aligned} F_{\mu\nu}^a &= \frac{-4}{g} \left[\frac{1}{x^4} (x^2 f'(x^2) - f(x^2) + f^2(x^2)) (\eta^a_{\nu\rho} x_\rho x_\mu - \eta^a_{\mu\rho} x_\rho x_\nu) \right. \\ &\quad \left. + \frac{1}{x^2} (-f(x^2) + f^2(x^2)) \eta^a_{\mu\nu} \right] \\ &= \frac{-4}{g} \left[\frac{x^2 f' - f(1-f)}{x^4} (\eta^a_{\nu\rho} x_\rho x_\mu - \eta^a_{\mu\rho} x_\rho x_\nu) - \frac{f(1-f)}{x^2} \eta^a_{\mu\nu} \right] \end{aligned}$$