

Analytic Properties of Scattering Amplitudes

Best to use a different form of the Radial Schrödinger equation: remove first derivatives using Riccati's transformation.

$$\left[-\frac{\hbar^2}{2m} \left(\frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r} - \frac{l(l+1)}{r^2} \right) + V(r) \right] R_l(r) = E R_l(r)$$

$$\left[\frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r} - \frac{l(l+1)}{r^2} + \underbrace{\frac{2m}{\hbar^2} E}_{k^2} - \underbrace{\frac{2m}{\hbar^2} V(r)}_{\equiv U(r)} \right] R_l(r) = 0$$

(wavenumber) Reduced potential

$$\left[\frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r} - \frac{l(l+1)}{r^2} + k^2 - U(r) \right] R_l(r) = 0.$$

Now, set $R_l(r) = \frac{1}{r} u_l(r)$ to obtain a new diff. eq. for $u_l(r)$:

$$\left[\frac{1}{r} \frac{\partial^2}{\partial r^2} - \frac{l(l+1)}{r^3} + \frac{1}{r} k^2 - \frac{1}{r} U(r) \right] u_l(r) = 0$$

$$\boxed{\left(\frac{\partial^2}{\partial r^2} - \frac{l(l+1)}{r^2} + k^2 - U(r) \right) u_l(r) = 0}$$

Note: If divided by k^2 , can express in terms of dimensionless quantity $\bar{r} = kr$.

All k -dependence is inside $\frac{U(r)}{k^2}$.

The free ($U(r)=0$) solutions are the Riccati-Bessel functions.

	<u>Riccati function</u>	<u>Small r behavior</u>	<u>asymptotic behaviour</u>
"sin"	$\hat{j}_l(kr) \equiv (kr) j_l(kr)$	$\frac{2^l l!}{(2l+1)!} (kr)^{l+1} + \dots$	$\approx \sin(kr - \frac{l\pi}{2})$
"cos"	$\hat{n}_l(kr) \equiv -(kr) n_l(kr)$	$\frac{(2l)!}{2^l l!} \frac{1}{(kr)^l} + \dots$	$\approx \cos(kr - \frac{l\pi}{2})$
"exp"	$\hat{h}_l^\pm(kr) \equiv \hat{n}_l(kr) \pm i \hat{j}_l(kr)$ $= \pm i (kr) h_l^{(\pm)}(kr)$	$\frac{(2l+1)!}{2^l l!} \frac{1}{(kr)^l} + \dots$	$\approx e^{\mp i \frac{l\pi}{2}} e^{\pm i kr}$ <small>(i)\mp(l)</small>

• Notice \hat{h}^\pm defined quite analogously to $\exp = \cos \pm i \sin$.

• \hat{j}_l is more regular at $r=0$ than \hat{n}_l is irregular.

Useful reflection identities:

$$\hat{j}_l(-kr) = (-1)^{l+1} \hat{j}_l(kr)$$

$$\hat{n}_l(-kr) = (-1)^l \hat{n}_l(kr)$$

$$\hat{h}_l^\pm(-kr) = (-1)^l \hat{h}_l^\mp(kr)$$

$$\hat{j}_l^*(z^*) = \hat{j}_l(z)$$

$$\hat{n}_l^*(z^*) = \hat{n}_l(z), \quad z \neq 0$$

$$[\hat{h}_l^\pm(z^*)]^* = \hat{h}_l^\mp(z), \quad z \neq 0$$

so that centrifugal barrier dominates.

To make rigorous statements about the analytic properties, make certain assumptions about $U(r)$:

Less singular than $1/r^2$ at origin: $\int_0 dr r U(r) \equiv \text{finite}$ (i.e. $U(r) \sim \frac{1}{r^{2-\epsilon}} + \dots$)

Vanishes faster than $1/r^3$ at infinity: $\int^\infty dr r^2 U(r) \equiv \text{finite}$ EXAMPLES: Yukawa, Woods-Saxon... but NOT Coulomb - does not fall off fast enough.

Potential analytic function of r for $\text{Re } r > 0$ "Analytic potentials"

As before, for each k , there are two linearly independent solutions. Different linear combinations defined by suitable boundary conditions yield solutions with useful analytic or physical properties:

(mixed boundary-value condition)
 ① PHYSICAL SOLUTIONS: $u_l(k, 0) = 0$ and $\lim_{r \rightarrow \infty} u_l(k, r) \rightarrow e^{i\delta_l(k)} \sin(kr - \frac{l\pi}{2} + \delta_l(k))$
 well-known & studied previously.

② REGULAR SOLUTIONS: $\phi_l(k, r)$ (Initial-value condition at $r \rightarrow 0$)
 $\lim_{r \rightarrow 0} \phi_l(k, r) \rightarrow \hat{j}_l(kr) \equiv \frac{2^l l!}{(2l+1)!} (kr)^{l+1} + O(r^{l+2})$
 — tends to physical, free radial solution at origin.

③ IRREGULAR SOLUTIONS: $f_l^\pm(k, r)$ (Initial-value condition at $r \rightarrow \infty$)
 $\lim_{r \rightarrow \infty} f_l^+(k, r) = \hat{h}_l^+(kr) \equiv e^{-i\frac{l\pi}{2}} e^{ikr}$
 $\lim_{r \rightarrow \infty} f_l^-(k, r) = \hat{h}_l^-(kr) \equiv e^{i\frac{l\pi}{2}} e^{-ikr}$
 — tend to free incoming and outgoing waves at infinity.